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The Magnetic Storm of December 16–17, 1917, as Recorded at Kew and Eskdalemuir Observatories.

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§ 1. In a recent paper,* I described a magnetic storm occurring on August 22, 1916, in which the disturbances recorded at Kew and Eskdalemuir Observatories were closely similar in type. The magnetic storm on December 16-17, 1917, affords in many ways a remarkable contrast, the changes recorded at the two observatories differing notably in type. The disturbance was the largest of the year. In the Kew declination (D) and horizontal force (H) curves, a prominent feature was a succession of oscillations with periods of about 20 minutes, and it was the hope of identifying these with corresponding oscillations at Eskdalemuir that suggested the present investigation. Dr. Crichton Mitchell, the Superintendent of Eskdalemuir Observatory, kindly sent the original curves, so the comparison was made under favourable conditions.

One of the obstacles to the identification of corresponding movements at the two stations is that, while H and D are recorded at Kew, the north (N) and west (W) components are recorded at Eskdalemuir. On instituting a minute comparison between the Kew H and D curves, it was found that, while oscillations with periods not far from 20 minutes were prominent in both, the times of the turning points by no means always agreed. A turning point in the H curve, for instance, might be represented by a short arrest or slight temporary reversal of movement in the D curve, the original direction of the D movement being almost immediately resumed. Again, towards the end of an oscillation, the trace might be nearly level for some minutes, or there might be a small short period oscillation, so that more than one choice was possible when assigning a time for the end of one major oscillation or the beginning of the next. Thus, when one had to compare Kew H with Eskdalemuir N curves, and Kew D with Eskdalemuir W curves, the fact that corresponding movements were difficult to identify is not surprising. Both observatories record vertical force (V), but the Kew V curves are so much disturbed by artificial electrical currents that they cannot be relied on for minor details. There were also difficulties special to the particular storm. Some of the N and W movements at Eskdalemuir were so rapid that the photographic trace was in places almost invisible. While the time marks

^{* &#}x27;Roy. Soc. Proc.,' A, vol. 93, p. 177.

occur on the base lines as well as on the curve lines, there is sensible parallax, and when the marks on the curve lines become invisible, the time cannot be fixed with quite the usual accuracy.

§ 2. The disturbance was preceded by a fairly quiet time, so its commencement may be accepted without hesitation as occurring between 8 h. and 9 h. on December 16. But the subsidence was as usual more gradual. Measurements were made of the curves from 8 h. on the 16th to 8 h. on the 17th. They were taken at four-minute intervals, because four minutes answers very nearly to 1 mm. in the base line. The curves were not smoothed, and the arithmetic mean of the measurements made throughout each hour G.M.T. was accepted as the hourly mean. Hourly values were thus obtained from 8.5 h. to 23.5 h. (11½ P.M.) on December 16, and from 0.5 h. to 7.5 h. on December 17. The arithmetic mean of these 24 hourly values was accepted as the mean for the "day" which commenced at 8 h. on December 16.

m: C 16 m		Kew.			Eskdalemui	r. ,
Time, G.M.T.	н.	D.	v.	N.	w.	v.
Dec. 16 8.5 h. 9.5 10.5 11.5 12.5 13.5 14.5 15.5 16.5 17.5 18.5 19.5 20.5 21.5 22.5 23.5 Dec. 17 0.5 1.5 4.5 5.5 6.5 7.5	7 + 9 - 6 + 10 + 34 + 12 + 38 + 75 + 70 + 131 + 100 + 71 - 15 - 36 - 175 - 100 - 21 + 11 + 68 + 68 + 4 + 18 + 8 - 1	7 - 8 - 2 - 8 - 10 - 13 - 2 + 6 - 34 - 54 - 63 - 73 - 118 - 139 - 117 - 65 - 70 - 65 - 69 - 51 - 59 - 40 - 43 - 31 - 28	7 - 8 - 13 - 13 - 15 + 6 + 14 + 22 + 48 + 114 + 174 + 184 + 182 + 144 + 82 + 59 + 80 + 61 + 35 - 8 - 19 + 69 + 91 + 98	7 + 9 - 18 - 12 - 11 - 4 + 1 + 15 - 22 + 28 + 132 + 122 - 2 - 117 - 156 - 63 - 55 - 79 - 153 - 95 - 21 - 33 - 21 - 22	7 - 3 - 7 + 4 + 20 + 27 + 64 + 121 + 118 + 108 - 65 - 252 - 123 - 43 - 14 - 27 - 0 - 27 - 15 - 3 - 8 - 13	7 - 6 - 10 - 13 - 14 - 5 - 5 + 194 + 297 + 288 + 279 + 158 - 54 - 99 - 24 + 1 - 8 - 147 - 149 - 47 - 18 - 6 + 1

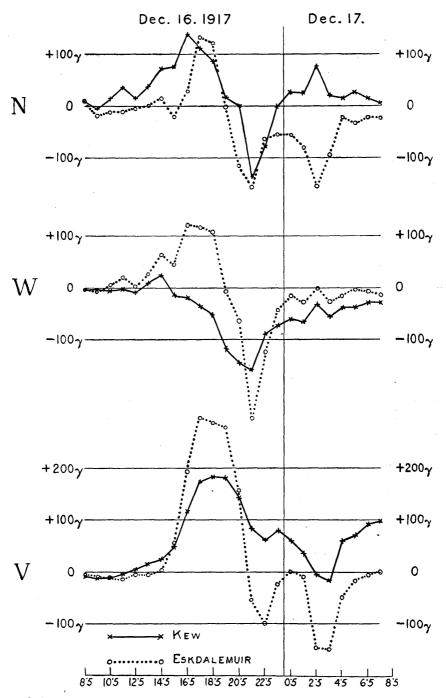
The Kew D trace records angular changes, but at present a rise of one minute in westerly declination is equivalent to a change of 5.36γ in force

directed 15° south of true west. The Eskdalemuir V trace unfortunately went off the sheet, in the direction of increasing force, several times between 16 h. and 19 h. on the 16th, there being no record for some 55 minutes in all, and the longest continuous absence of trace being about 20 minutes. When the curve was off the sheet, the value answering to the margin of the sheet was accepted. Judging by the trend of the curve, the margin was not much exceeded in most cases. But between 17 h. and 18 h., when the trace was off the sheet for 20 minutes at a time, the margin may have been considerably exceeded. An error x in one hourly mean implies an error x/24 in the mean for the day.

The existence of this uncertainty in the mean value of V at Eskdalemuir was one of the reasons which led to the adoption as standard values of quiet day means derived from the previous 24 hours. Table I shows the departure of the hourly means from these quiet day means, the unit being for all the elements $1 \gamma \equiv 1 \times 10^{-5}$ C.G.S. magnetic unit. Maximum and minimum values are in heavy type. Every hourly value of D at Kew, except that for 14.5 h., falls short of the quiet day mean, i.e., represents a position of the declination needle to the east of the normal. The existence of this phenomenon at Kew, while there is no corresponding abnormality in W at Eskdalemuir, is one of the most remarkable differences between the two stations. Another remarkable feature is that while N at Eskdalemuir shows the depression customary in that element and in H towards the end of a magnetic storm, there is no corresponding sensible depression in H at Kew. V was much enhanced at both places in the afternoon of the 16th, a customary phenomenon of magnetic storms. This elevation of V was followed by a rapid fall before midnight, also a customary feature; but subsequent to this there was another considerable rise, and a second rapid fall, more especially at Eskdalemuir.

§ 3. Hourly mean values of N and W at Kew were calculated from the hourly means of H and D, and the departures of these from the quiet day means are shown in the accompanying diagram, with the corresponding departures in N and W at Eskdalemuir and in V. The scale of ordinates is the same for the six curves.

Up to 22 h. on December 16 there is a general resemblance between the changes in N at the two stations; but later, especially from 1 h. to 3 h. on the 17th, the changes are more nearly diametrically opposite. The N ranges for the whole 24 hours, as given by the hourly means, were 288 γ at Eskdalemuir as compared with 279 γ at Kew. This gives a very inadequate idea of the excess in the activity at Eskdalemuir. A better idea is given by the values of the absolute range, *i.e.*, the excess of the largest over the least value



recorded at any instant during the 24 hours. These absolute ranges were 620 γ at Eskdalemuir in N, and 408 γ at Kew in H.

The resemblance between the W diagrams for Kew and Eskdalemuir is not close, but it is not less close than usual between 1 h. and 3 h. on the 17th, when the N diagrams from the two stations were opposed. At Esk-dalemuir W was markedly enhanced from 13 h. to 19 h. on the 16th, whereas at Kew W was markedly depressed after 15 h. The ranges from the hourly means were 373γ at Eskdalemuir and 183γ at Kew. The difference between the absolute W range at Eskdalemuir and the absolute D range at Kew was even greater, the former being 588γ , the latter only 189γ .

The ranges from the V hourly means were 446 γ for Eskdalemuir and 197 γ for Kew, the corresponding absolute ranges being respectively 552 γ and 259 γ . As already explained, the values for Eskdalemuir are necessarily under-estimates.

§ 4. Table II gives the results of measurements of the oscillations of longer period already referred to. The different curves were considered and measured quite independently. A minimum reading was taken as the commencement of each oscillation. The second minimum for one oscillation served in many cases as the first minimum for the next. In other cases an interval was left, answering to a portion of the curve which was level or showed one or more short period oscillations, the inclusion of which in either of the two adjacent longer period oscillations would have been arbitrary. If we started with the assurance that we had to do with a continuous train of oscillations, confused by the superposition of waves of shorter period, we should naturally assign the doubtful piece of curve to one or other of the adjacent longer period oscillations. If there had been a very pronounced resemblance between Kew and Eskdalemuir curves, and the hypothesis of a small differential time error had markedly improved the coincidence of turning points, it might have been fairly justified. But, as matters stood, it seemed best to leave the figures exactly as the curve measurements made them.

The result is that 14 longer period oscillations were recognised in both the H and D curves at Kew, the first 13 of which proceeded with little or no interruption, whilst the 14th followed after an hour's interval. The 13th differed somewhat in type from the others. It possibly represented two oscillations rather than one. The mean duration for the first 12 oscillations, accepting the figures in Table II, is in H 20·4 minutes, and in D 21·7 minutes. If we assume that there was a continuous train of oscillations, and accept the times given for the beginning of the 1st and the end of the 12th oscillation, the mean durations become 21·8 minutes for H and 22·5 minutes for D.

At Eskdalemuir, between 15 h. and 20½ h., on the 16th, 16 oscillations of longer period were recognised in N and 13 in W. Some, however, were of

Table II.—Oscillations of Longer Period.

	Fall.	7. 442. 331. 332. 331. 113. 113. 446. 614. 614.		Fall.	7. 38. 38. 37. 192 192 192 193 193 102
				<u> </u>	
	Rise.	286 286 288 288 288 288 388 388 388 388 388 388		Rise	7. 683 683 70 1134 1150 1150 1150 1150 1150 1150 1150 115
	Interval.	Hin. 24. 22. 22. 23. 23. 24. 11. 11. 11. 11. 11. 12. 23. 23. 23. 23. 23. 23. 23. 23. 23. 2		Interval.	Hin. 23 24 25 24 25 25 25 26 26 26 26 26 26 26 26 26 26 26 26 26
Kew D.	2nd min.	h. m. 15 36 16 18 16 18 16 18 17 23 17 23 17 23 17 23 17 23 18 21 19 50 19 15 19 15 19 15 19 15	Eskdalemuir W	2nd min.	2nd min. 15 36 15 36 16 18 16 11 16 23 16 23 17 4 17 4 17 4 18 8 18 24 19 46 20 25
	Max.	h. m. 15 28 16 39 16 5 16 28 16 28 17 17 17 40 17 51 18 38 19 58 19 52 21 18	Es	Max.	h. 15 31 15 31 16 31 16 6 16 18 17 17 17 17 17 17 17 17 18 35 19 34 20 8
	1st min.	h. m. 15 12 15 36 16 56 16 22 16 22 17 3 17 23 17 23 17 23 17 23 18 21 18 50 19 20 19 20 19 42		1st min.	h. m. 15 13 15 13 16 11 16 11 16 11 17 17 17 18 18 8 18 19 18 19 18 19 18
	Fall.			Fall.	7. 86 90 90 225 33 34 35 1153 1153 105 105 105 143 143
	Rise.	22 52 52 52 52 53 53 54 55 56 56 56 56 57 57 57 57 57 57 57 57 57 57 57 57 57		Rise.	644 644 869 869 861 861 1443 1443 1443 1443 1443 1443 1443 14
	Interval.	min. 144 22 24 25 25 25 25 25 25 25 25 25 25 25 25 25		Interval.	min. 231 23 21 38 10 10 10 10 10 10 10 10 10 10 10 10 10
Kew H.	2nd min.	h. m. 15 36 16 24 16 24 17 7 17 7 17 80 17 80 18 26 18 26 19 22 20 25 20 25 20 25	Eskdalemuir N.	2nd min.	b. 15 832 H. 16 832 H. 17 4 833 H. 17 4 834 H. 18 18 18 18 18 18 18 18 18 18 18 18 18
	Max.	h. m. 15 28 15 28 16 50 16 40 16 40 17 17 17 17 17 55 18 18 18 18 58 19 28 20 9	Ä	Max.	b. 28 m. 15 28 m. 16 28 m. 16 28 m. 16 28 m. 17 17 17 28 m. 18 34 4 4 19 34 4 55 m. 19
	1st min.	h. m. 115 22 116 36 116 0 116 24 117 48 117 48 118 8 118 29 119 22 119 44		1st min.	h. h. 15 11. 15 32. 16 20 16 28 16 44 17 19 17 19 18 18 18 18 18 19 15 1
, A		1	į		1 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2

decidedly shorter period than the others, and may have answered to portions only of the oscillations recognised at Kew, large short-period oscillations occurring sometimes at Eskdalemuir while the Kew curve showed almost stationary conditions.

The Eskdalemuir W oscillations could be brought into fair parallelism with the H and D oscillations at Kew by assuming that Nos. 3 and 4 at Eskdalemuir together answered to No. 3 at Kew, whilst Nos. 5–11 at Eskdalemuir corresponded respectively to Nos. 4–10 at Kew, Nos. 12 and 13 at Eskdalemuir corresponding to the oscillations similarly numbered at Kew. We could also bring the earlier N oscillations at Eskdalemuir into fair parallelism with the others by regarding the pairs Nos. 3 and 4, 5 and 6, 7 and 8, and 9 and 10 as each but a single oscillation, answering respectively to Nos. 3, 4, 5, and 6 at Kew.

The Eskdalemuir curves contained some exceedingly large rapid oscillations, practically unrepresented at Kew. These must presumably have arisen from some source of disturbance which was either purely local, or was situated very much nearer to Eskdalemuir than to Kew.

As regards the size of the oscillations recorded in Table II, the mean values of the rise and fall at Kew were respectively $55\,\gamma$ and $60\,\gamma$ for H and $27\,\gamma$ and $35\,\gamma$ for D. The H movements were thus, on the average, decidedly the larger. If we accept the arithmetic mean of the rise and fall as a measure of the amplitude, we find that the first six oscillations in D were very similar in amplitude, while the corresponding H oscillations differed widely. There was, in short, little, if any, parallelism between the amplitudes of the oscillations in the two elements. If we accept the problematical scheme of correspondence between Kew and Eskdalemuir oscillations given above, we find that the Eskdalemuir ranges were almost invariably the larger, but the numerical results, from a comparison of Kew and Eskdalemuir curves, are too uncertain to be worth recording.

§ 5. During many of the rises and falls recorded in Table II, the rate of change of the element varied greatly. Rates of change derived from Table II would thus give an inadequate idea of the rapidity of the more rapid changes. An independent set of measurements was accordingly made, the results of which appear in Tables III-VII. They give the central time of the movement, *i.e.* the arithmetic mean of the times of beginning and ending, the number of minutes the movement lasted, and the estimated rapidity of the change of force in γ 's per minute. In most cases, where the rises and falls recorded in Table II showed a sudden marked acceleration or retardation, measurements were made on shorter portions of curve having a nearly uniform slope. In a few cases, however, mean rates are given for

longer periods of time, when there was a conspicuous movement in one direction, interrupted by only trifling arrests or reversals. No measurements were attempted of the shorter period movements, mostly small, occupying one or two minutes or less, unless they were quite exceptionally large. The slope of the curve then affords a ready check on the result derived from the time measurements.

Table III.—Rapid Changes in H at Kew.

Mean time.	Dura- tion.	Change.	Rate.	Mean time.	Dura- tion.	Change.	Rate.	Mean time.	Dura- tion.	Change.	Rate.
h. m. 15 8.5 25 0 32 0 43 0 53 0 16 2 5 12 0 21 5 27 5 43 0 53 5 17 3 0 12 0 18 5 26 5	min. 3 6 8 14 6 5 2 5 7 2 11 8 10 3 7	732 +22 -14 +50 -21 +49 +23 -48 +35 -35 +59 -94 +81 -444	γ/min. 11 3 · 5 1 · 8 3 · 5 10 11 9 · 5 5 · 0 17 5 · 5 12 8 · 0 15 8 · 0	h. m. 17 38 0 45 0 51 5 18 0 5 9 5 19 0 34 0 44 0 54 0 19 4 5 15 5 24 0 38 0 48 5	min. 8 6 7 11 3 14 10 10 4 5 13 4 12 9	7. +60 -24 +27 -28 +18 -83 +77 -65 +21 -32 -63 +32 -77 +53	γ/min. 7 * 5 4 * 0 2 * 5 6 * 0 6 * 0 7 * 5 5 * 0 6 * 5 5 * 0 6 * 5 6 * 0	h. m. 20 1·0 11·0 22·5 21 9·0 15·5 21·0 27·0 34·0 54·0 22·32·5 0 59·0 1 3·0 2 14·0 55·5	min. 10 4 5 8 5 8 4 2 10 39 2 6 20 29	7. + 57 - 37 - 41 - 215 + 83 - 27 - 46 + 25 + 46 + 89 - 25 + 99 - 111	γ/min. 5 · 5 · 5 9 · 0 8 · 0 27 17 4 · 5 11 12 4 · 5 2 · 3 14 4 · 0 5 · 0 4 · 0

Table IV.—Rapid Changes in D at Kew.

Mean time.	Dura- tion.	Change.	Rate.	Mean time.	Dura- tion.	Change.	Rate.	Mean time.	Dura- tion.	Change.	Rate.
h. m. 15 8 5 20 0 31 0 37 5 50 0 16 0 5 15 0 25 0 31 5 43 0 48 0 17 0 0	min. 3 16 6 3 12 9 6 6 7 2 8 6	723 +19 -40 +19 -28 +26 -30 +26 -14 +28 -25	$\gamma/\text{min.}$ 7 · 5 1 · 2 · 6 · 5 6 · 5 · 5 2 · 3 · 0 5 · 0 4 · 5 · 5 7 · 0 3 · 5 4 · 0	h. m. 17 8 · 0 20 · 0 35 · 0 43 · 0 48 · 5 18 15 · 5 27 · 0 38 · 0 19 0 · 5 11 · 5 23 · 0 34 · 0	min. 10 6 10 6 5 11 12 10 5 7 6 16	7. +20 -34 +48 -22 +17 -42 +37 -14 -22 +13 -37	γ/min. 2·0 5·5 5·0 3·5 4·0 4·0 3·5 3·0 3·0 2·2 2·3	h. m. 19 47 0 20 6 5 21 6 0 16 0 21 0 32 5 49 0 0 59 0 1 3 0 2 41 0 58 5	min. 10 13 8 4 6 5 28 2 6 12 23	7. + 33 - 32 - 35 + 54 - 50 + 44 + 76 + 18 - 21 + 40 - 44	γ/min. 3 · 5 2 · 5 4 · 5 13 8 · 5 9 · 0 2 · 5 9 · 0 4 · 0 3 · 5 1 · 9

Table V.—Rapid Changes in N at Eskdalemuir.

Mean time.	Dura-	Change.	Rate.	Mean time.	Dura- tion.	Change.	Rate.	Mean time.	Dura-	Change.	Rate.
h. m. 15 9 0 13 0 20 5 29 5 36 0 49 5 58 5 16 5 5 9 0 14 0 22 0 25 0 35 5 38 0 49 0 56 5 59 5 17 2 5 6 0 9 5 12 5 23 5 26 5 32 0	min. 4 11 5 8 11 7 3 4 6 4 2 3 2 10 3 3 3 3 3 6	7. - 51 + 26 + 38 - 86 + 64 - 89 + 87 - 25 + 61 - 39 + 35 + 40 - 22 - 30 + 109 - 153 + 149 - 161 - 161 - 55	7/min. 13 6 5 3 5 17 8 0 8 0 12 8 5 15 6 5 8 7 13 11 15 11 13 12 50 38 12 15 54 9 0	h. m. 17 41·0 47·5 52·0 18 8·5 26·0 33·5 36·5 39·5 43·0 47·5 19 4·0 18·5 29·5 37·0 45·5 9·0 12·5 21·5 33·5 38·5 58·5	min. 2 5 4 5 10 1 5 1 4 5 22 7 3 6 7 4 5 2 3 5 5 5 3	7 30 + 64 - 64 - 85 + 217 + 32 - 102 + 29 + 46 - 38 - 217 + 35 - 79 - 84 + 36 - 30 - 30 + 33 - 52 + 25 - 30 + 29	γ/min. 15 16 17 22 32 20 29 11 7 · 5 10 · 0 9 · 0 12 13 12 4 · 0 12 15 10 5 · 0 6 · 0 9 · 5	h. m. 21 19·5 31·0 43·5 51·0 22 0·0 8·5 12·0 24·0 51·5 56·5 23 12·0 0 0·0 51·5 56·5 20·5 36·5 50·5 50·5 58·0 3 3·5 8·0 48·0	min. 15 8 7 6 2 5 2 8 7 3 2 4 4 2 4 3 3 5 5 5 7 8 3 6 20	7282 +193 + 92 + 53 - 28 - 30 + 25 - 45 - 33 + 24 + 28 - 37 - 45 - 31 - 49 - 30 + 69 - 30 + 60 - 36 + 42 + 84	y/min. 19 24 13 9 14 6 12 5 5 5 4 5 8 0 14 7 18 11 10 16 14 6 12 11 7 5 12 7 0 4 0

Table VI.—Rapid Changes in W at Eskdalemuir.

									,		
Mean time.	Dura- tion.	Change.	Rate.	Mean time.	Dura- tion.	Change.	Rate.	Mean time.	Dura- tion.	Change.	Rate.
h. m. 14 36 0 15 8 5 27 0 33 5 38 0 6 2 5 7 0 14 5 20 5 24 0 35 5 46 0 49 5 56 6 17 4 0 12 0 24 0 31 0 38 0 45 0	min. 2 5 4 7 7 2 7 5 2 3 1 2 5 3 6 10 4 6 8 6	7. + 20 - 59 + 31 - 38 + 32 - 80 + 128 - 35 + 80 - 121 + 34 + 38 - 48 - 36 + 88 + 54 - 167 + 203 - 99 - 58 + 100 - 43	γ/min. 10 12 7 · 5 8 · 0 11 18 17 11 24 17 13 48 18 18 18 18 28 20 25 9 · 5 12 7 · 0	h. m. 17 51 0 18 9 5 18 0 29 0 38 0 43 0 19 4 0 15 0 25 5 32 0 43 5 50 0 55 0 55 0 12 5 24 0 12 5 24 0 12 5 24 0 12 5 24 0 12 5 24 0 12 5 25 5 20 6 12 5 21 4 5 8 5 12 0	min. 6 3 8 10 4 2 6 6 6 3 4 4 5 5 8 2 5 4 4 3 2 3 5 3 4 4	7. + 52 + 22 - 58 + 84 - 31 + 26 - 68 - 66 + 40 + 27 - 492 + 58 - 27 + 58 - 27 + 58 - 27 + 58 - 27 + 23 - 58 - 42 - 178	γ/min. 8 · 5 7 · 0 8 · 5 7 · 5 13 11 13 6 · 5 12 18 7 · 0 13 11 5 · 5 19 17 7 · 5 50 47	h. m. 21 16 5 23 · 0 30 · 0 37 · 0 38 · 5 22 5 · 0 17 · 5 23 57 · 5 1 0 · 5 22 · 5 21 7 · 0 30 · 5 38 · 0 41 · 0 50 · 5 59 · 0 12 · 0 20 · 5 35 · 5	min. 5 8 6 2 1 9 6 7 3 3 5 9 12 3 2 4 7 4 6 8 5 15	7. +224 -171 + 59 + 54 - 35 + 59 + 42 + 36 + 22 + 25 - 32 + 83 - 24 + 28 - 32 - 69 + 22 - 29 + 27 - 42	γ/min. 45 21 10 27 35 6.5 7.0 5.0 7.5 8.5 6.5 7.0 8.0 14 8.0 10 5.5 5.0 3.5 5.0 3.5 5.0

Mean time.	Dura- tion.	Change.	Rate.	Mean time.	Dura- tion.	Change.	Rate.	Mean time.	Dura- tion.	Change.	Rate.
h. m. 15 33 · 5 54 · 0 16 1 · 0 6 · 5 20 · 0 25 · 5 43 · 0 48 · 0 51 · 5	min. 9 8 6 5 8 3 6 4 3	γ. + 35 + 36 - 28 + 32 + 130 - 50 + 35 + 56 - 43	γ/min. 4 · 0 4 · 5 4 · 5 6 · 5 16 17 6 · 0 14	h. m. 18 8·0 19 43·0 50·0 20 2·5 11·5 19·0 32·5 21 11·0 15·5	min. 8 4 10 9 6 15 2	7 32 + 17 - 74 - 37 + 26 - 32 - 56 - 89 - 229	γ/min. 4 · 0 · 4 · 0 · 7 · 5 · 4 · 0 · 3 · 0 · 5 · 5 · 3 · 5 · 44 · 46	h. m. 21 25 0 35 0 38 5 50 5 22 1 5 21 0 2 25 5 3 12 5 42 5	min. 2 2 5 9 5 10 55 25 35	7. + 22 - 19 + 47 - 30 - 43 + 38 - 223 + 91 + 87	γ/min. 11 9·5 9·5 3·5 8·5 4·0 4·0 3·5 2·5

Table VII.—Rapid Changes in V at Eskdalemuir.

Independent measurements were first made of the absolute times of the beginning and end of each movement. A differential measurement was then made giving the duration directly. Finally, the slope of the curve was measured with a special scale. With times measured only to the nearest minute, no very great accuracy can be claimed for the rates, but the checks mentioned above should be a sufficient protection against large errors. The probable error is naturally greatest where the time interval is least, and the large rates obtained for some of the shortest movements may thus incur suspicion. The Kew rates are all much slower than the faster Eskdalemuir rates, and the fastest of them, 27 y per minute in H, was derived from an 8-minute interval, and so can hardly be much in error. The fastest rate for N changes at Eskdalemuir, viz. 54 y per minute, is based on a 3-minute interval, but that interval was immediately followed by a second 3-minute interval, the change occurring during which had the same estimated rapidity. The two movements constituted an oscillation in which the to-and-fro movements were practically equal. The duration of the complete oscillation was thus a particularly convenient one to measure. If we accept six minutes for it, but suppose that an error was made in allotting equal times to the . two movements, then the duration of one of them must have been overestimated, and so its rapidity under-estimated. A somewhat similar argument applies in the case of the next most rapid N rate, viz., 50 γ per minute.

The most rapid rate for W changes in Table VI, viz. 50 γ per minute, is followed by rates of 47 γ , 44 γ , and 45 γ per minute, and these four rates arise from intervals which form between them the 17 consecutive minutes 21 h. 2 m. to 21 h. 19 m. on the 16th. The error in the total interval is unlikely to have exceeded one minute, and any redistribution of the time

amongst the four sub-intervals would almost inevitably have led to at least one rate in excess of any of the four assigned.

The fastest rate for V arose from a 5-minute interval. It was a particularly favourable case for measuring the slope of the curve, as the trace was bold and the slope uniform to the eye. In the case of the longer intervals, the assigned rate must in general have been very sensibly exceeded during part of the time.

No V movement was included in Table VII, any part of which was lost through the trace going off the sheet. Several rapid movements were excluded for this reason, but none so rapid as the fastest movement given in the Table.

No rates were calculated for V changes at Kew, because little reliance could have been placed on them.

		Rises.			Falls.		Ri	Rises and falls.			
Element.	Aggregates.		3.5	Aggregates.		36	Aggregates.		Mean		
	Dura- tion.	Move- ment.	Mean rate.	Dura- tion.	Move- ment.	Mean rate.	Dura- tion.	Move- ment.	rate.		
Kew H Kew D Eskdalemuir N Eskdalemuir W Eskdalemuir V	mins. 188 146 183 178 137	γ. 1034 529 2214 2069 775	γ/min. 5·5 3·6 12·1 11·6 5·7	mins. 175 158 196 156 138	γ. 1217 576 2430 2163 985	7·0 3·6 12·4 13·9 7·1	mins. 363 304 379 334 275	γ. 2251 1105 4644 4232 1760	γ/min 6·2 3·6 12·3 12·7 6·4		

Table VIII.—Aggregates of Rapid Changes and Corresponding Mean Rates.

Table VIII gives particulars of the aggregates of the rises and falls included in Tables III–VII. But for the loss of trace, the time aggregate for V at Eskdalemuir would have been considerably larger. While the time aggregates for the different elements are not identical, we may fairly conclude that during the most active part of the storm changes in N and W at Eskdalemuir were roughly twice as rapid as changes in H, and $3\frac{1}{2}$ times as rapid as changes in D at Kew. The V changes at Eskdalemuir and the H changes at Kew had very similar mean rates, but the fastest V changes at Eskdalemuir were decidedly faster than the fastest H rates at Kew. The V traces were, however, of a very different type from the others, oscillations of short period being much less in evidence. Considering that the ranges of the diurnal inequalities in N and W at Eskdalemuir in December are only about 15γ and 20γ respectively, it is a little startling to find that during

six hours of one particular "day," the mean rates of change per minute in these elements were quite 12γ . The contributions to the aggregates of changes from the other 18 hours of the 24 would have been considerable, and the mean rate of change in N and W for the whole 24 hours could hardly have been less than 4γ per minute.

§ 6. Any complete estimate of the expenditure of energy during a magnetic storm is probably impossible. A magnetic storm is usually accompanied by aurora and earth currents. In this particular case, according to observations made by Prof. Störmer near Christiania, the aurora extended from 100 to 400 kiloms above the earth's surface. There is no known method of finding the intensity at any height of the electric currents producing aurora. Few stations attempt to measure earth currents, and little, if anything, is known as to the depth to which earth currents extend. Thus the energy expended in aurora, and the energy expended in earth currents, are alike inaccessible to calculation. If the ultimate source of all the terrestrial phenomena is some form of electrical discharge from the sun, the energy represented by the various phenomena in the earth and the atmosphere is probably but an insignificant fraction of the whole.

The estimate of the energy represented by the changes shown in magnetic curves is naturally a simpler problem, but even it is fraught with difficulties, practical as well as theoretical. For what has been done to give definiteness to the problem we are mainly indebted to the late Prof. Bidlingmaier.*

The formula given by Maxwell for the energy in a magnetic field when there are no electric currents is

$$(1/8\pi)\iiint (\alpha^2+\beta^2+\gamma^2)\,dx\,dy\,dz,$$

where α , β , γ represent the rectangular components of magnetic force. The integral is supposed to be taken throughout the whole of the magnetic field. It takes, moreover, as point of departure a total absence of force. In the present case we know the absolute values of the three components at a fixed point, and the sequence of changes in these components; but the intensity of the field never vanishes, the changes encountered being in fact small compared with the mean values of the elements. Suppose we confine our attention for the moment to one of the three components. The value may be regarded as given by the curve ordinate y at the instant. If y_0 be the accepted normal value, the "activity," as Bidlingmaier has called it, at the instant is

$$(1/8\pi)(y-y_0)^2. (1)$$

^{* &#}x27;Veröffentlichungen des Kaiserlichen Observatoriums in Wilhelmshaven,' "Ergebnisse der Magnetischen Beobachtungen im Jahre 1911," Neue Folge, Heft 2.

What we are concerned with in practice is the mean value of the "activity" for some specified time, whether an hour, a day, or a year. Let

$$y = y_h + \eta, \tag{2}$$

where y_h is the mean value for the hour under consideration. Then, from the properties of the arithmetic mean, we have

Mean value of $(y-y_0)^2$ per hour = $(y_h-y_0)^2$

+ mean value of
$$\eta^2$$
 per hour. (3)

To determine the mean value of η^2 , Bidlingmaier measured the curve at six-minute intervals and took $(1/10) \Sigma \eta^2$ as the mean value for the hour. In the present case the curve was measured at four-minute intervals, viz., at 0 4, ..., 56, 60 minutes after the hour, and

$$(1/15) \Sigma \eta^2 \equiv \{\frac{1}{2} (\eta_0^2 + \eta_{15}^2) + \eta_1^2 + \dots + \eta_{14}^2\},\tag{4}$$

was accepted as the mean. Thus $(1/8\pi)(1/15)\Sigma\eta^2$ answers to what Bidlingmaier called A_h^x for the particular hour. The mean "activity" for the one hour is

$$A_h^x + (1/8\pi)(y_h - y_0)^2. (5)$$

Suppose now we want the mean activity for the day. If y_d be the mean value for the day, and y_1, \ldots, y_{24} the mean values for the several hours, then

Mean value of
$$(y_h - y_0)^2 = (y_d - y_0)^2 + (1/24) \{ (y_1 - y_d)^2 + \dots + (y_{24} - y_d)^2 \}.$$
 (6)

In Bidlingmaier's notation

$$(1/8\pi)\left\{(y_1 - y_d)^2 + \dots + (y_{24} - y_d)^2\right\} \equiv A_d^h,\tag{7}$$

$$(1/8\pi)(y_d - y_0)^2 \equiv A_a{}^d. \tag{8}$$

His complete expression for the mean "activity" of the day is

$$A_h^x + A_d^h + A_a^d. (9)$$

In this connection A_h^x represents the arithmetic mean of the 24 individual hourly values of A_h^x . The curve ordinates are to be regarded as expressed in terms of $1\gamma (\equiv 1 \times 10^{-5}$ C.G.S.), and the unit in the "activity" results is 1×10^{-10} erg per cubic centimetre.

A difficulty at once arises as to A_a^d , viz., what to accept for the normal value y_0 . Bidlingmaier apparently thought that theoretically the best plan would be to derive it from the mean value for the whole year, by assuming the rate of secular change uniform. For instance, for Wilhelmshaven in 1911, taking $+16\gamma$ per annum as the rate of change of H, he accepted as the value of y_0 for a day n days subsequent to July 1

Mean for year
$$+(n/365)16\gamma$$
.

His value of y_0 thus increased gradually throughout the year.

The mean value for a year is unknown until the year is completed. For the secular change it is desirable to know the mean value for the subsequent as well as the previous year. Thus the calculation of $A_a{}^d$ in the way outlined above would have to be postponed to a somewhat indefinite date. Recognising this, though he gave values of $A_a{}^d$ at Wilhelmshaven for all days from January 1 to June 30, 1911, the period to which he confined himself, Bidlingmaier concluded that in general the only practical course would be to omit $A_a{}^d$, i.e., to accept for y_0 the mean value of the individual day. This would give for the day's mean "activity" $A_b{}^x + A_d{}^b$, the symbols having the same meaning as before.

In accepting as the best theoretical value for y_0 one based on the mean value for the year, Bidlingmaier recognised one difficulty. The mean value is not the same when we confine ourselves to quiet days as when we include all days of the year. This is especially true of H. The quiet day yearly mean would seem the more natural one to employ in the present connection, but that is a point on which opinions may differ.

Another difficulty is that an annual inequality is believed to exist in the case of all the magnetic elements. If an annual inequality really exists, it should presumably be taken into account in the present problem. These considerations support the view that the practical difficulties in the way of applying to individual days a correction $A_a{}^d$ based on the mean value of the element for the year and the secular change are very serious. When we examine the results for individual days of 1911, which Bidlingmaier found at Wilhelmshaven, we find that $A_a{}^d$ is by no means trifling. The following selection from Bidlingmaier's results should carry conviction on this point:—

	D alone.				H alone.	•		D+H.		
	A_h^x .	\mathbf{A}_{d}^{h} .	$A_a{}^d$.	A_h^x .	A_{a}^{h} .	$\Lambda_a{}^d$.	A_h^x .	\mathbf{A}_{a}^{h} .	$\mathbf{A}_a{}^d$.	
Mean from 6 months April 9, 1911 March 20, 1911		9·72 53·8 37·5	1:15 48:8 0:0	1 · 73 14 · 9 14 · 6	5 ·40 35 ·6 53 ·5	2.05 38.2 2.0	4 ·28 30 ·4 32 ·0	15 ·12 89 · 4 91 · 0	3 ·20 87 ·0 2 ·0	

On the average day the value of $A_a{}^d$ is here about three-quarters that of $A_h{}^z$, and on some individual days $A_a{}^d$ is of the same order as $A_d{}^h$. Our view as to the relative "activities" on April 9 and March 20, two of the most disturbed days of 1911, for instance, would be largely determined by the retention or neglect of $A_a{}^d$.

The adoption of Bidlingmaier's "activities" as an international scheme, in place of the present "character" figures, was under consideration before

the war. The question whether a correction of the type of $A_a{}^d$ should be applied ought, I think, to receive very careful consideration before any final decision is reached.

If the present method of selecting the international quiet days continues, perhaps the most satisfactory course would be to accept the mean value from the five quiet days as the normal for all days of the month. At present three months are dealt with together, and the announcement of the quiet days selected, even in pre-war times, did not appear until several months had elapsed, after the end of the quarter. No doubt matters could be accelerated by dealing with the months individually. If Bidlingmaier's scheme were substituted for the present scheme, the selection of the quiet days would entail the pre-existence of the "activity" statistics. In that event, possibly, the best course might be to accept for y_0 the mean value of the last previous day considered quiet at the particular observatory.

§ 7. In the case of a disturbance so large as that of December 16-17, 1917, a difference of 2γ or 3γ in y_0 is of minor importance. I decided to calculate two sets of "activity" results. The first set accepts for y_0 a mean derived from the "day" ending at 8 h. on December 16. The second set adopts Bidlingmaier's suggestion, and neglects $A_a{}^d$ entirely, accepting for y_0 the mean value for the "day" commencing at 8 h. on December 16. For the disturbed day the mean of the 24 hourly means was taken. For the previous or quiet day, the mean was taken from measurements made at 14 h. and 20 h. on the 16th, and at 2 h. and 8 h. on the 17th. The mean of readings taken at these four hours on the average quiet day agrees very closely with the mean from the 24 hours in all the elements. With a view to possible future intercomparisons it is desirable to put on record the two sets of mean curve ordinates.

		Kew.		F	Eskdalemuir.			
	н.	D.	v.	N.	w.	v.		
Day of storm Quiet day	mm. 46 ·2 44 ·0	30 ·4 39 ·4	mm. 73·2 69·4	mm. 16 ·8 21 ·8	mm. 14 ·9 15 ·7	mm. 61·5 55·1		
Excess of disturbed day mean Force equivalent	+ 2 ·2 + 13γ	-9·0 -48γ	+3.8 +617	-5 ·0 -26γ	-0.8 -4γ	+6·4 +28γ		

At Eskdalemuir the equivalents of 1 mm. of ordinate were 5.34γ in N, 4.96γ in W and 4.33γ in V. At Kew the scale value determinations gave

as the equivalents of 1 mm. of ordinate 5.9γ in H, 5.36γ in D, and 16.0γ in V.

In presenting the results obtained from the use of the quiet day mean, I have not followed Bidlingmaier's method of obtaining a correction $A_a{}^d$. This is convenient when we confine ourselves to the mean "activity" for the whole day, but seems less appropriate when we consider individual hours. In fact it is not altogether clear how Bidlingmaier proposed to treat individual hours, except when $A_a{}^d$ is neglected. In his Table XIII,* when dealing with the diurnal variation of "activity," he takes no cognizance of $A_a{}^d$. He gives, however, two sets of figures. The first set includes mean values obtained for each hour of the day from the five international quiet days of the month, the second set gives the excess over these values of the corresponding values obtained when all days of the month are included. The second set of figures are regarded by Bidlingmaier as giving the diurnal variation of disturbance. They would obviously be unaffected by any correction which was common to the all-day and quiet-day "activities."

I am doubtful whether Bidlingmaier's method of deducing a diurnal variation of disturbance is satisfactory. The differences between the curve ordinates at one and the same hour of the five quiet days of a month must represent in the main disturbance, thus it is only what is common to the five days that can well claim to represent entirely quiet conditions.

In the present case, even if I had entirely approved Bidlingmaier's procedure, it would have been impracticable to adopt it, because the selection of quiet days for December, 1917, is unlikely to be announced until midsummer, 1918.

§8. The best course seemed to be to refer each hourly measurement *directly* to the quiet day mean accepted, so that the complete mean value of the "activity" for the hour becomes

$$A_h^x + (1/8\pi)(y_h - y_0)^2. \tag{10}$$

Tables IX and X give the results thus calculated for individual hours at Kew and Eskdalemuir. The unit employed in the Tables is that used by Bidlingmaier, viz., 1×10^{-10} erg per cubic centimetre.

Bidlingmaier's notation being complicated, the following notation has been employed in the Tables and their discussion

$$A_{2} \equiv (1/8\pi)(1/15) \Sigma \eta^{2}
A_{1} = (1/8\pi)(y_{h} - y_{0})^{2}
A_{1}' = (1/8\pi)(y_{h} - y_{d})^{2}$$
(11)

^{*} Loc. cit., p. 29.

Here η has the same meaning as in (2), while y_0 and y_h are the mean values for the days which end and commence respectively at 8 h. on December 16.

The excess in the mean value of A_1 for the day over that of A_1' is equivalent to Bidlingmaier's $A_a{}^d$. The values derived for the excess from (8), accepting the differences between y_d and y_0 given in § 7, are as follows, in terms of the unit employed in the Tables:

All the preliminary arithmetical operations which went to the construction of the Tables were performed on differences of curve ordinates expressed in millimetres or minutes of arc, the unit 1×10^{-10} erg being introduced only in the final operation; thus an absolute identity was not to be expected between the differences between A_1 and A_1 just calculated, and those derived in Tables IX and X by taking the actual means of the 24 hourly values.

It may be well to add that a calculation was made to ascertain the order of magnitude of the corrections that would have been applied if an attempt had been made to allow for the "activity" present on quiet days. The calculations were applied to the *mean* diurnal inequalities for December as given by five selected quiet days a month.

The Kew data were for the 11 years 1890–1900 combined, the Esk-dalemuir data for the single year 1914. The means of the 24 hourly values thus found for A_1' (Bidlingmaier's A_d^h) were in the usual unit:

```
At Kew, for H, 0.28; for D, 1.02; for V, 0.10; total for 3 elements 1.40; At Eskdalemuir, for N, 0.45; for W, 0.66, for V, 0.07; total for 3 elements 1.18.
```

No data were available for the calculation of A_2 (or A_h^x), but the contributions from A_2 on a quiet December day must be very small. If the quiet days had been treated individually, larger mean values would have been found for A_1 . But, if we may judge from Bidlingmaier's data for January, 1911,* the whole corrections with the quiet days treated individually would not have exceeded double the figures given above, and the largest correction for any individual hour would have been only from three to four times the mean. The effect of any such corrections in Tables IX and X would have been insignificant.

The calculations were carried to the first decimal place; but it has been omitted in all the values of the total "activities," as well as in all individual cases where the value of A_2 , A_1 , or A_1 exceeded 100. Results are given for the two horizontal components combined, as well as for the sum from the

Table IX.—" Activity" at Kew from 8 h. December 16 to 8 h. December 17, 1917. (Unit 1×10^{-10} erg per cubic centimetre.)

	$\Lambda_2 + \Lambda_1'$.	260 230 295 257	170 201 345 183 704 863	817 874 736 1791 574 93	285 8 285 8 285 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8	392
	A_2+A_1 , A_2+A_1'	11 18 23 60	18 73 263 372 1354	1817 1938 1683 2209 759 481	328 379 175 280 280 412 415	637
H+D+V	Α ₁ ′.	257 321 286 249	165 194 327 144 668	762 819 700 1615 520 79.7	11.7 46.5 311 263 5.9 49.8 78.3	363
_	Α ₁ .	8 3 13 3 51 0	13.9 65.5 245 333 1318 1763	1762 1883 1648 2033 705 468	321 240 290 154 203 276 371 414	809
	A ₂ .	60 0 0 0 0 4 6 0	47 1888 3988 3988 3988 3988 3988 3988 3988	55.2 35.8 176 53.8 13.8	741 888 144 685 748 874 861 861	29.7
	$A_2+A_1. A_2+A_1'.$	69 111 76 82	52 115 285 162 570 343	199 276 457 1756 573	130 190 100 100 104 147 137	528
	A_2+A_1 .	7 10 15 56	16 64 241 266 813 590	453 613 853 1924 619 225	175 197 357 147 68 89 89 41	328
H+D.	Α,'.	67.6 102 67.3 75.1	48 ·8 269 137 556 309	159 226 426 1598 520 65·3	11.7 121.5 121.6 121.6 5.0 5.0 13.9 23.7	206
	Λ_1 .	5.8 1.6 6.5 50.0	12.5 57.7 226 241 800 557	413 564 821 1766 567 213	173 191 288 139 64.4 64.4 86.6	305
	A ₂ .	1.6 8.6 4.6	24.3 1.5 1.5 1.3 1.3 1.3 1.3 1.3 1.3 1.3 1.3	40 · 3 49 · 7 31 · 9 158 52 · 3 11 · 7	0.00 0.00 0.00 0.00 0.00 0.00 0.00 0.0	23.3
	A_1' .	190 218 218 174	116 84·3 57·5 6·7 112 518	603 593 274 17·6 0·2 14·4	0 8 8 8 8 9 9 9 9 9 9 9 9 9 9 9 9 9 9 9	157
۷.	Α ₁ .	25.5 6.7 1.0	1.4 7.8 19.3 91.8 518 1206	1348 1320 826 268 139 255	148 483 1139 139 330 382 382	303
	A ₂ .	1.00 8.00 9.00	112 122 123 123 123 20	14.9 8.9 9.5 1.1 7.1	001 001 001 001 001 001	6.4
	Α,'.	67.0 88.0 67.0 57.5	48:8 84:3 116 7:8 7:8	24.9 195 330 190 11.5 19.3	111 5 17 6 0 4 4 8 4 8 8 8 8 9 10 0 1 1 1 2 0 9 1 1 5 0	57 .6
Ď.	Α,	ស ០៧4 ចំពុំចំំំ	6.7 0.2 1.4 46.1 116 158	212 555 770 545 168 196	168 190 104 139 63.7 73.7 38.3 31.2	150
	A ₂ .	0200	1.4 0.6 2.9 10.4 4.7 10.3	10.6 6.0 22.3 49.1 7.7 0.9	0.71 17.8 0.2 0.1 0.1 0.1	9.9
	Α1′.	0.6 14.4 0.4 17.6	0.0 24.9 153 129 555 301	134 31·2 95·7 1408 509 46·1	0.2 121 3.2 3.2 3.2 1.0	148
H.	A ₁ .	3.5. 4.1. 4.0. 1.94	5.7 224 195 684 398	201 9.0 51.6 1220 398 17.6	4.1 184 0.0 0.0 0.0 0.0 0.0 0.0	155
	A ₃ .	27.46 10.00	125 8 8 8 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4	29.7 43.7 29.6 109 44.6 10.8	11.1 44.4 511.6 5.0 0.3 0.3 4.0	17.6
Hour.	G.M.1.	$\begin{array}{c} 8-9\\ 9-10\\ 10-11\\ 11-12\\ \end{array}$	12-13 $13-14$ $14-15$ $15-16$ $16-17$ $17-18$	18-19 19-20 20-21 21-22 22-28 23-38	010004700F	Mean for "day"

Table X.—" Activity" at Eskdalemuir from 8 h. December 16 to 8 h. December 17, 1917. (Unit 1×10^{-10} erg per cubic centimetre.)

N+W+V.	$A_2 + A_1'$	101 84 95 110	67 118 304 2093 4753	4334 2753 1264 4070 1385 222	68 209 2169 1623 235 89 49 49	1102
	A2+A1. A2+A1'	α 4 4 8 8	37 195 313 2366 5012	4631 3303 1821 4324 1285 239	132 305 2099 1444 126 60 22 22 27	1161
	A ₁ ′.	97.8 60.4 74.8 104	61.5 111 281 129 1841 4504	4066 2553 1151 3362 1251 200	64 ·3 185 1863 1452 224 86 ·3 86 ·3 32 ·9	892
	Α ₁ .	5.0 18.8 13.1 28.6	1.8 30.1 172 224 2114 4763	4362 3103 1708 3616 1151	128 280 1794 1273 115 56 7 21 4 26 1	1051
N+W.	Ą,	24.0 204.0 304.0 306.0	252 252 252 252 252 252 252	268 200 112 708 134 22·0	3.8 3.06 5.47 171 110 0.0 0.0 0.9	110
	A2+A1. A2+A1'.	45 22 23 40 40	26 75 276 150 850 1846	1633 199 538 3563 128	39 122 255 11 11 5 2	472
	A ₂ +A ₁ .	88 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5	6 36 194 153 721 1495	1318 178 774 3968 870 205	132 101 123 144 125 127 128 127 127 128 127 127 127 127 127 127 127 127 127 127	501
	Α ₁ ′.	51.8 2.9 11.0	20.7 67.3 255 100 743 1600	1372 23·1 478 3094 619 91·8	35 2 133 2 203 2 5 3 3 1 5 5	400
	A _l .	3.6 14.9 6.4 20.8	0.8 29.1 172 104 615 1249	1058 1.6 714 3500 761 194	128 278 933 389 265 201 260	429
	Α2.	23.8 20.5 6.0	5.4 7.3 22.0 49.5 107 246	260 176 60 0 469 109 10 6	2021 2021 2022 2022 2022 2022 2022 2022	72.4
	A ₁ ′.	46.1 57.5 63.7 70.3	40 ·8 43 ·4 26 ·9 29 ·0 1098 2904	2693 2530 673 268 632 108	29.0 51.6 1220 1248 218 84.3 46.1 29.0	592
Α.	A_1 .	4.0.5 7.9 8.7	1.0 1.0 0.2 121 1499 3514	3305 3101 995 116 390 22 •9	0.0 861 885 880 112.9 0.0	622
	A ₂ .	0000	0.2 0.0 145 39.0 33.1	80 22 52 52 53 54 74 74	0.5 227 119 5.5 1.0 0.4	37.8
	A ₁ '.	00 8 42 24 22 6	38 · 3 184 100 623 593	500 0·2 148 2431 564 60·6	4.0 0.6 0.6 19.3 4.8 0.0 0.0	222
W.	Α ₁ .	4.0 0.0 0.0 0.0	0 · 2 28 · 0 163 84 · 3 555	465 1.4 168 2530 603 73.7	r 2002 8000 8000 8000 8000 8000	223
	Α2.	0 4 1 4 6 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	2.6 5.0 14:3 18:7 41:6 115	24.2 86.1 143 15.0 145 15.1 8 5.1	1884 4041 100 100 100 100 100 100 100 100 1	29.5
	Α ₁ ′.	51.6 2.5 7.8 9.0	19.3 29.0 70.3 0.6 121 1007	873 222 9 330 663 54 5	31.2 112 643 184 1.0 2.0 1.0 0.6	178
z.	Α ₁ .	3.2 12.9 5.7 4.8	0.0 0.0 19.3 31.2	593 0·2 545 970 1158	121 249 933 360 17.6 43.4 17.6 19.3	205
	A ₂ .	2.0 19.0 19.5 1.4	22.8 7.7 30.8 64.9	236 90.1 326 64.3 64.3	11.1 34.0 42.0 0.0 0.0 0.0 0.0 0.0	45.9
Hour.	G.M.T.	$\begin{array}{c} 8-9.\\ 9-10\\ 10-11\\ 11-12 \end{array}$	12–13 13–14 14–15 15–16 16–17	18—19 19—20 20—21 22—22 23—23	0 1 2 6 4 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6	Mean for "day"

three components, because it was the only way of intercomparing the Kew and Eskdalemuir results from the horizontal plane, and because Bidlingmaier omitted vertical force.

The relative importance of A_2 and A_1' is of special interest in connection with the question of the calculation of A_2 , a very laborious process when carried out strictly according to the formula. If we express A_2 as a percentage of $A_2 + A_1'$ in the case of the mean values for the day, we obtain the following results:—

```
At Kew, in H, 11; in D, 9; in V, 4; in H+D, 10; in H+D+V, 8; At Eskdalemuir, in N, 21; in W, 11; in V, 6; in N+W, 15; in N+W+V, 10.
```

With these we may compare Bidlingmaier's mean results at Wilhelmshaven for January to June, 1911; viz., in H, 24; in D, 21; in H+D, 22. It would appear, however, that most of Bidlingmaier's results were derived not from actual measurement of $(1/10) \Sigma \eta^2$, but from measurements of the hourly ranges, and certain numerical relationships which he believed to hold between the value of A_h^z and the square of the corresponding hourly range. In a recent paper,* I arrived at the conclusion that Bidlingmaier's relationships were not generally satisfactory. If they had been accepted in the present case, we should unquestionably have got a higher mean value for A_2 . If $A_2/(A_2+A_1')$ were always as small as the above values for Kew suggest, it could not be claimed that any very high degree of accuracy is essential in the method of calculating A_2 .

Bidlingmaier found that on the average day the "activity," whether A_2 , A_1 , or A_1 , was considerably greater for D than for H. On the present occasion, the "activity" at Kew, especially A_2 and A_1 , is conspicuously larger for H than for D. The A_2 and A_1 results for N and W at Eskdalemuir differ less, which suggests that the horizontal direction in which the changes of force were most active made a fairly close approach to the magnetic meridian.

In temperate latitudes the range of the regular diurnal inequality in V is considerably less than in either horizontal component. Thus, undoubtedly, on the average day at any European station, the contribution from V will be small compared with the sums of the contributions from the horizontal components. But, on the present occasion, the contribution from V to A_2+A_1 or A_2+A_1' at Kew is nearly equal to the combined contributions from H and D, while at Eskdalemuir the contribution from V considerably exceeds that from N and W combined. This shows that the question of the

^{* &#}x27;Terrestrial Magnetism and Atmospheric Electricity,' vol. 22, p. 57 (1917).

exclusion of V data from an international scheme possesses an importance which is not suggested by a study confined to ordinary days.

§ 9. It is interesting to note, confining ourselves to the results for H+D that the mean value of A_2+A_1' at Kew on December 16-17 was 11.8 times that of the average day from January 1 to June 30, 1911, at Wilhelmshaven, and 15 times that of the average January day. Also, the mean value of A_2+A_1' at Kew between 21 h. and 22 h. on December 16 was 52 times that of the corresponding hour of the average January day at Wilhelmshaven. It is obvious that, when we confine ourselves to data from a single year, a few highly disturbed days might exercise a swamping influence in statistics as to the annual or diurnal variation of magnetic "activity."

During the earlier part of the storm the "activity" at Kew was not inferior to that at Eskdalemuir, but after 16 h. on the 16th the Eskdalemuir "activity" was much the greater. At both stations, so far as the horizontal field is concerned, 21 h.-22 h. on the 16th was decidedly the most disturbed hour. This was the time at which aurora was brightest, according to reports from meteorological stations, several Scottish stations reporting a corona not far from the zenith. Kew and Eskdalemuir also agree in showing a maximum of "activity" in V between 17 h. and 20 h. on the 16th, and a marked recrudescence of "activity" between 2 h. and 4 h. on the 17th, which was presumably associated with auroral streamers observed at Southport between $2\frac{1}{4}$ h. and $2\frac{3}{4}$ h.

§ 10. In view of the great labour involved in applying Bidlingmaier's scheme, and the various uncertainties, I suggested in 'Terrestrial Magnetism' (loc. cit.) that a much simpler scheme, if applied at the numerous observatories which participate in the present international scheme, might suffice for the magnetic classification of days. The suggestion was that, if R_h , R_d , R_v represent the absolute ranges in the three magnetic elements, $R_h^2 + R_d^2 + R_v^2$, or, if the horizontal field only is considered, $R_h^2 + R_d^2$ might be accepted as a measure of the mean "activity" for the day. If the horizontal components recorded were N and W, then $R_n^2 + R_{w^2}$ would take the place of $R_h^2 + R_d^2$.

I have compared the mean monthly values of the "activity" at Wilhelmshaven for three of the six months dealt with by Bidlingmaier, with the corresponding arithmetic means of the squares of the Kew daily ranges. The following were the results obtained:—

	January.	March.	May.	Mean.
$\frac{\text{Mean value of } \Lambda_2 + \Lambda_1' \text{ at Wilhelmshaven}}{\text{Mean value of } R_{\lambda}^2 + R_{\alpha}^2 \text{ at Kew}}$	0.000196	0 .000224	0 .000194	0 • 000205
$\frac{\text{Mean value of } A_2 + A_1 \text{ at Wilhelmshaven}}{\text{Mean value of } R_k^2 + R_{k^2} \text{ at Kew}}$	0.000221	0 .000251	0.000235	0 .000236

Taking now the "activities" and ranges found for December 16-17, we get the following results:—

At Kew-

$$\begin{aligned} &(\text{Mean A}_2 + \text{A}_1' \text{ for H} + \text{D})/(\text{R}_h^2 + \text{R}_d^2) &= 0.000113, \\ &(\text{Mean A}_2 + \text{A}_1' \text{ for H} + \text{D} + \text{V})/(\text{R}_h^2 + \text{R}_d^2 + \text{R}_v^2) &= 0.000146, \\ &(\text{Mean A}_2 + \text{A}_1 \text{ for H} + \text{D})/(\text{R}_h^2 + \text{R}_d^2) &= 0.000162, \\ &(\text{Mean A}_2 + \text{A}_1 \text{ for H} + \text{D} + \text{V})/(\text{R}_h^2 + \text{R}_d^2 + \text{R}_v^2) &= 0.000237. \end{aligned}$$
 At Eskdalemuir—
$$&(\text{Mean A}_2 + \text{A}_1' \text{ for N} + \text{W})/(\text{R}_n^2 + \text{R}_w^2) &= 0.000065, \end{aligned}$$

 $\begin{aligned} &(\text{Mean A}_2 + \mathbf{A}_1' \text{ for N + W})/(\mathbf{R}_n^2 + \mathbf{R}_w^2) &= 0.000065, \\ &(\text{Mean A}_2 + \mathbf{A}_1' \text{ for N + W + V})/(\mathbf{R}_n^2 + \mathbf{R}_w^2 + \mathbf{R}_v^2) &= 0.000106, \\ &(\text{Mean A}_2 + \mathbf{A}_1 \text{ for N + W})/(\mathbf{R}_n^2 + \mathbf{R}_w^2) &= 0.000069, \\ &(\text{Mean A}_2 + \mathbf{A}_1 \text{ for N + W + V})/(\mathbf{R}_n^2 + \mathbf{R}_w^2 + \mathbf{R}_v^2) &= 0.000112. \end{aligned}$

The difference between the above figures for the two stations arises from A_1 and A_1 . If we confined ourselves to A_2 we should get in place of the above ratios: For the horizontal components, 0.000012 at Kew, 0.000010 at Eskdalemuir; for all three components, 0.000011 at both places.

The means of the absolute daily ranges for the whole year 1914 were

At Kew,
$$44\gamma$$
 in H, and 55γ in D;

At Eskdalemuir, 60y in N, and 58y in W.

This makes

$$(R_n^2 + R_w^2)/(R_h^2 + R_d^2) = 1.40.$$

Wilhelmshaven being very nearly midway as regards latitude between Kew and Eskdalemuir, the absolute ranges at Wilhelmshaven probably exceed those at Kew. Thus if we had employed Wilhelmshaven instead of Kew ranges for comparison with the Wilhelmshaven "activities" of 1911, we should most likely have got smaller values than those obtained above for the ratios.

The scheme which I suggested does not postulate that the ratio borne by "activities" to squares of ranges should be the same for all stations. It might well be different at a quiet station like Helwan, and at a disturbed

station like Sitka. Everything considered, however, it is probable that on any scheme which assumed "activity" proportional to the square of the absolute range, the mean "activity" for December 16-17 would have been sensibly over-estimated at Kew, and considerably over-estimated at Eskdalemuir. On the other hand, the large difference between the values of A_1 and A_1 in Table IX, and the uncertainty this implies, shows that even with the elaborate procedure entailed by Bidlingmaier's full scheme, there is a large probable error at individual stations.

Under these circumstances a simple scheme, even if admittedly imperfect, deserves consideration.

A Method of Avoiding Collision at Sea. By J. Joly, Sc.D., F.R.S., a Commissioner of Irish Lights.

PART I. (Received May 8, 1918.)

The following method of avoiding collision at sea depends on the use of synchronised signals, transmitted in different media. Such signals, travelling at different rates, enable the distance of their source to be inferred by observation of the gain in time of the faster upon the slower travelling signal. Thus, if signals be simultaneously emitted by wireless and by submarine bell (or Fessenden oscillator), the former being transmitted with practically infinite velocity, the latter arrive with a lag which is the time the submarine sound requires to traverse the intervening medium. The rate of propagation of sound in water being closely 4800 feet per second, the lag is 0.62 second for one-half sea-mile.

In practice the signals may be so ordered as to dispense with the stop-watch or chronograph. This is accomplished by sending out the wireless ticks in groups of, say, 20 "dots" spaced to intervals of 0.6 second. The stroke of the bell precedes the first of these dots by one of these intervals. Thus, when the sailor is half mile from the source he hears the first wireless dot along with the bell stroke. If he is 1 mile distant the bell stroke comes in with the second dot, and so on. He has, in fact, only to count up the dots till he hears the bell, and the number of the dot coincident with the bell is the number of half sea-miles intervening between his ship and the source of the signals. It is possible to estimate the quarter mile by noting a want